

E-cloud map formalism: an analytical expression for quadratic coefficient

T. Demma, INFN-LNF, Frascati (Italy),
S. Petracca, A. Stabile, University of Sannio, Benevento (Italy) & INFN Salerno, Italy.

Abstract

The evolution of the electron density during electron cloud formation can be reproduced using a bunch-to-bunch iterative map formalism. The reliability of this formalism has been proved for RHIC [1] and LHC [2]. The linear coefficient has a good theoretical framework, while quadratic coefficient has been proved only by fitting the results of compute-intensive electron cloud simulations. In this communication we derive an analytic expression for the quadratic map coefficient. The comparison of the theoretical estimate with the simulations results shows a good agreement for a wide range of bunch population.

INTRODUCTION

In [1] it has been shown that, the evolution of the electron cloud density can be described introducing a quadratic map of the form:

$$n_{m+1} = \alpha n_m + \beta n_m^2 \quad (1)$$

where n_{m+1} and n_m are the average densities of electrons between two successive bunches. The coefficients α and β are extrapolated from simulations and are functions of the beam parameters and of the beam pipe characteristics. An analytic expression for the linear map coefficient that describes electron cloud behavior from first principles has been derived for straight sections of RHIC [3]. In this paper we find an analytical expression the quadratic term coefficient. We consider $N_{el,m}$ quasi-stationary electrons gaussian-like distributed in the transverse cross-section of the beam pipe. The bunch $m+1$ accelerates the $N_{el,m}$ electrons initially at rest to an energy \mathcal{E}_g . After the first electrons-wall collision two new jets are created: the backscattered electrons with energy \mathcal{E}_g and the "true secondaries" (with energy $\mathcal{E}_0 \sim 5 \text{ eV}$).

The sum of these jets gives the number of surviving electrons $N_{el,m+1}$, then one gets the linear coefficient

$$\alpha = \frac{N_{el,m+1}}{N_{el,m}} \quad (2)$$

In the next section we compute the quadratic term coefficient β when the saturation condition of the electron cloud is obtained. Once calculated saturation we pass to estimate theoretically the coefficient β . We compare our results with the outcomes of numerical simulations obtained using ECLLOUD [4]. In the Table 1 we report all parameters used for our calculations.

Table 1: Input parameters for analytical estimate and ECLLOUD simulations.

Parameter	Unit	Value
Beam pipe radius b	m	.045
Beam size a	m	.002
Bunch spacing s_b	m	1.2
Bunch length h	m	.013
Energy for δ_{max} $\mathcal{E}_{0,max}$	eV	300
Energy width for secondary e^-	eV	-
Number of particles per bunch N_b	10^{10}	$4 \div 9$
Secondary emission yield (max) δ_{max}	-	1.7
Secondary emission yield ($\mathcal{E} \rightarrow 0$)	-	.5

STEADY-STATE: ELECTRONIC DENSITY OF SATURATION

In the chamber we have two groups of electrons belonging to cloud: primary photo-electrons generated by the synchrotron radiation photons and secondary electrons generated by the beam induced multi-pactoring. Electrons in the first group generated at the beam pipe wall interact with the parent bunch and are accelerated to the velocity given by: $v/c = 2\bar{N}_b r_e / b$, where r_e is the classical electron radius and \bar{N}_b is the effective value of bunch population and

$$\bar{N}_b = \frac{h}{h + s_b} N_b \quad (3)$$

s_b being the bunch spacing and h the length of bunch. Electrons in the second group, generally, miss the parent bunch and move from the beam pipe wall with the velocity given by: $v/c = \sqrt{2\mathcal{E}_0/mc^2}$, \mathcal{E}_0 being the average energy of the secondary electrons, until the next bunch arrives. The process of the cloud formation depends, respectively, on two parameters:

$$k = \frac{2\bar{N}_b r_e h}{b^2} \quad (4)$$

$$\xi = \frac{h}{b} \sqrt{\frac{2\mathcal{E}_0}{mc^2}} \quad (5)$$

The second one is the distance (in units of b) passed by electrons of each group before the next bunch arrives. At low currents, $k \ll 1$, each electron interacts with many bunches before it reaches the opposite wall. In the opposite extreme case, $k > 2$, all electrons go wall to wall in one bunch spacing. The transition to the second regime occurs when $k \sim 1$. The density of the secondary electrons grows until the space-charge potential energy of the

secondary electrons is lower than \mathcal{E}_0 . The saturation condition can be obtained by requiring that the potential barrier is greater than electron energy in the point $r/b = 1 - \xi$

$$-eV(1 - \xi) \sim \mathcal{E}_0 \quad (6)$$

where V is the electric potential generated by the bunch and the electron cloud. To calculate the electric potential we assume that our system is composed by a chamber with radius b , a bunch with radius a and length h , an electron cloud with density ρ . We consider the following electron distribution :

$$\rho(r) = \rho_0 e^{-\frac{(r - r_0)^2}{2\sigma^2}} \quad (7)$$

where ρ_0 is fixed by the condition

$$2\pi h \int_a^b \rho(r) r dr = -N_{el} e \quad (8)$$

and N_{el} is the total number of electrons in the volume $\pi h(b^2 - a^2)$. The electric potential $V(r)$, defined by the condition $V(b) = 0$ is:

$$V(r) = -V_0 \left[g \ln x + \frac{G(x)}{F(1)} \right], \quad (9)$$

where $F(x) = \int_{\tilde{a}}^x \exp(-(\tilde{y} - \tilde{r}_0)^2 / 2\tilde{\sigma}^2) y dy$, $G(x) = \int_x^1 F(y)/y dy$, $g = \bar{N}_b / N_{el}$, $V_0 = N_{el} e / 2\pi\epsilon_0 h$ and $x = r/b$, $\tilde{a} = a/b$, $\tilde{r}_0 = r_0/b$, $\tilde{\sigma} = \sigma/b$. We note that if $\sigma \gg b$ (or $\tilde{\sigma} \gg 1$) and $r_0 = 0$ we obtain the uniform electron cloud and with $a \rightarrow 0$ we must neglect the radial dimension of bunch with respect to that one of electron cloud. In this case equation (9) gives

$$V(r) = -V_0 \left[g \ln x + \frac{1 - x^2}{2} \right] \quad (10)$$

Obviously the potentials depend on g , the ratio of the densities of the beam and of the cloud averaged over the beam pipe cross-section. In FIG. 1 we report the spatial behavior of two potentials. The potential (10) has minimum at $r = r_m = b\sqrt{g}$ and is monotonic for $g > 1$ within the beam pipe. For $g < 1$ it has minimum at the distance $r_m < b$, and the condition $g = 1$ defines the maximum density. this is the well known condition of the neutrality. The condition formulated in this form is, actually, independent of the form of distribution. Similar behavior is found also for the gaussian distribution density and is compared with respect to previous one (FIG. 1). By imposing the condition (6) we find the critical number (saturation condition) of electrons in the chamber

$$N_{el,sat} = \frac{2\pi\epsilon_0 h F(1) \mathcal{E}_0}{e^2 G(1 - \xi)} - \frac{F(1) \ln(1 - \xi)}{G(1 - \xi)} \bar{N}_b \quad (11)$$

while the average density of saturation is found by assuming that electrons are confined in a cylindrical shell with

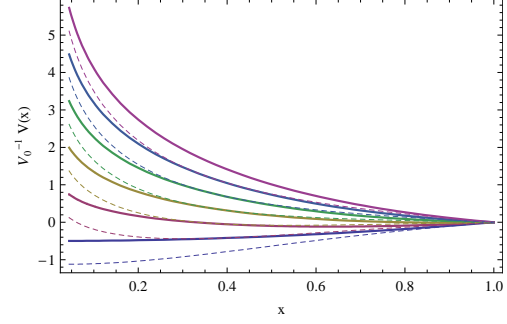


Figure 1: Plot of $V_0^{-1} V(x)$, (9) and (10)), in the case of uniform (solid lines) and gaussian (dashed lines) electronic distribution for $g = 0 \div 2$, $\tilde{a} = .04$, $\tilde{r}_0 = 0$, $\tilde{\sigma} = .3$.

inner radius a and external radius $r_0 + p\sigma$ where p is a free parameter. So

$$n_{sat} = \frac{N_{el,sat}}{\pi h b^2 [(\tilde{r}_0 + p\tilde{\sigma})^2 - \tilde{a}^2]} \quad (12)$$

where p is a free parameter. For a uniform electron cloud distribution we find the saturation density

$$\bar{n}_{sat} = \frac{\bar{N}_{el,sat}}{\pi h b^2 [1 - \tilde{a}^2]} \quad (13)$$

In the FIG. 2 we show the behavior of saturation density (12) and (13). It is obvious for a gaussian distribution we get a estimate of density saturation greater than that of a uniform distribution. In fact, the same number of electrons occupies a smaller volume (due to the Gaussian distribution).

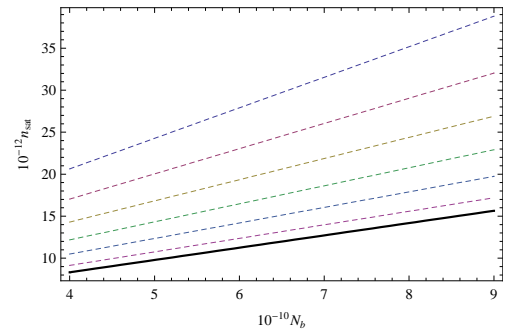


Figure 2: Plot of electronic densities of saturation n_{sat} vs N_b , (12) and (13)), withf uniform (solid line) and gaussian (dashed lines) electronic distribution for $\tilde{a} = 0.04$, $\tilde{r}_0 = 0$, $\tilde{\sigma} = 0.3$ and $p = 2 \div 3$.

ANALYTICAL DETERMINATION OF COEFFICIENTS

The coefficient β can be found by imposing the saturation condition of map (1):

$$n_{sat} = \alpha n_{sat} + \beta n_{sat}^2 \rightarrow \beta = \frac{1 - \alpha}{n_{sat}} \quad (14)$$

and the map (1) becomes

$$n_{m+1} = \alpha n_m + \frac{1-\alpha}{n_{sat}} n_m^2 \quad (15)$$

In Fig. (3), (4) we show the trends of the coefficient (14) as a function of δ_{max} for various values of bunch population and viceversa.

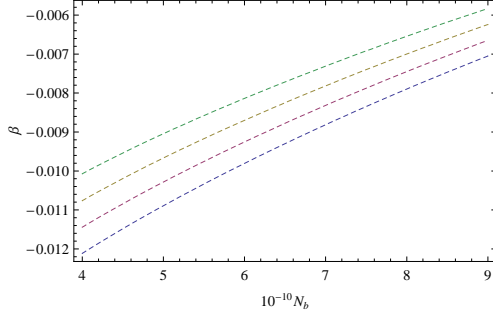


Figure 3: Analytical prediction of coefficient β (14) for values $\delta_{max} = 1.4 \div 2$ and $p = 2$.

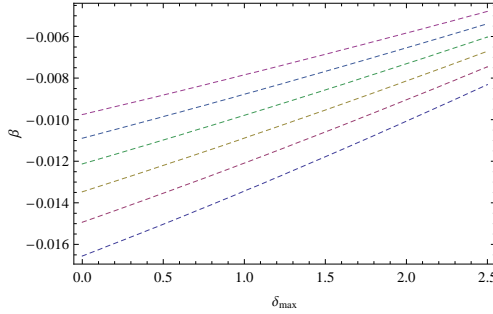


Figure 4: Analytical prediction of coefficient β (14) for values $N_b = 4 \div 9$ and $p = 2$.

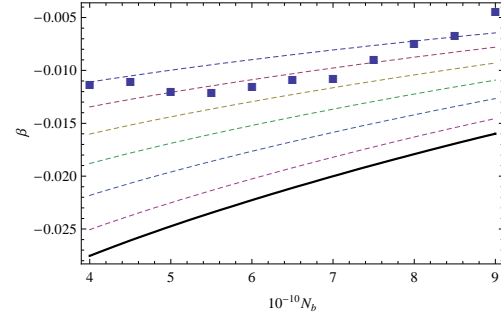


Figure 5: Comparison of the quadratic coefficient β (Eq. (14)) derived using ECLLOUD simulations (points) and using the analysis of previous sections (dashed lines) with $p = 2 \div 3$. The solid line is the result by assuming an uniform density.

RESULTS AND CONCLUSIONS

In Figs. 5 the analytical behavior and the outcomes of simulations (ECLLOUD code) of β coefficient using the parameters reported in Table 1 show an acceptable agreement. As a future work the analytical result could be useful to determine safe regions in parameter space where to minimize the electron clouds. Furthermore we would extend our results to include the presence of a magnetic field.

REFERENCES

- [1] U.Iriso and S.Peggs, "Maps for Electron Clouds", Phys.Rev. ST-AB8, 024403, 2005.
- [2] T.Demma et al., "Maps for Electron Clouds: Application To LHC", Phys.Rev.ST-AB10, 114401 (2007).
- [3] U. Iriso and S. Pegg. Proc. of EPAC06, pp. 357-359.
- [4] <http://wwwslap.cern.ch/electron-cloud/Programs/Ecloud/ecloud.html>.